MEG実験 液体キセノンガンマ線検出器の性能改善

OUTLINE

- Introduction
- New developments
 - face factor optimization
 - dead channel recovery
 - non-uniformity correction
 - alternative energy reconstruction

MEG EXPERIMENT



Target sensitivity : Br($\mu \rightarrow e\gamma$)~10⁻¹³



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N DETECTOR

- 900L liquid xenon
- 846 PMTs







ENERGY RESOLUTION

energy resolution key to sensitivity

 $R_{acc} \propto (R_{\mu})^2 * (\Delta \Theta)^2 * (\Delta E_{\nu})^2 * \Delta T * \Delta E_e$

• evaluated with 55MeV γ from π^0 runs $(\pi + p \rightarrow \pi^0 + n)$



- π^0 decay provides 55–83MeV γ ray
- Monochromatic γ obtained by selecting back-to-back opening angle

ENERGY RECONSTRUCTION

• weighted sum of scintillation photons

$$N = \sum Q_i / g_i / q e_i \times w_i$$

- Energy scale determined at 55MeV in π^0 run
- correction in position (non-uniformity) and change in light yield



Friday, September 16, 11

FACE FACTOR OPTIMIZATION

- Optimize weight factor of each face for each position
- Energy is reconstructed as a linear combination of each faces ($\sum w_i \times nface[i]$)
- Optimize by minimizing the variance in π^0 peak





FACE FACTOR OPTIMIZATION



Optimize weights at each position

Parameterize weights and optimize at the same time

Dead Channel Recovery

- 8 PMTs in 2010
- No dead channels on inner face
- Little effect on resolution







NON UNIFORMITY CORRECTION

- Correct position dependent detector response
- Calibration with CW-Li peak and π^0 -55MeV peak
- Depth correction with π^0 peak
- position correction with CW-Li peak
- Included additional position correction with π^0 -55MeV peak
- Remaining non-uniformity ~0.2%

- Cockcroft-Walton accelerator
 - Nuclear reaction by protons
 - Li(p, γ)Be 14.6,17.6MeV
 - B(p, γ)C 4.4, 11.7MeV





Depth correction functions for different position



Alternative Energy Reconstruction

- Treat each PMT as a detector and fit energy for each PMT
- Calculate event energy from tube energies
 - weighted mean energy of truncated distribution
- Correction table for gain × QE is prepared from output of PMT for each event
- Potential advantages:
 - can optimize PMT selection
 - insensitive to vast variations in solid angles subtended by nearby PMTs

ENERGY SCALETIME VARIANCE

- Light yield monitored with several calibration sources(CW 18MeV γ, Ni 9MeV γ, AmBe 4.4MeV γ, cosmic ray peak)
- While each has good stability(0.2%), they show the same trend
- Time-dependent correction by combination of calibrations



COSMIC RAY REJECTION

- Cosmic rays that enter from inner face and stop in LXe volume can be accidental background. •
- Cut on ratio of inner and outer charge •
 - Rejection optimized at 1% inefficiency (from MC signal) •
- Introduced additional cut using waveform (inefficiency<0.1%) ٠



PMT Sum 30%

Raw data Nov-Ave Conts

Template Fit



OPENING ANGLE EFFECT



180

83

Eγ

• We choose back to back events in π^0 runs

- Estimate opening-angle effect with data & MC, and unfold the contribution
- Actual resolution is better than the measured by ~0.15%

PERFORMANCE

- resolution before: ~2%
- current resolution: 1.9%



SUMMARY

- Energy resolution improved by introducing several new corrections
 - face factor optimization
 - non-uniformity
 - light yield correction
 - CR rejection
 - Opening angle effect
- Alternative energy reconstruction method studied
- Future improvements possible