



Photon Detector

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1. Large Prototype Study

- ✍ Gamma beam test
- ✍ Xenon purification
- ✍ Absorption length measurements

2. Other related R&D

- ✍ PMT development
- ✍ Purity monitor

3. Final Photon Detector

- ✍ Expected performance (MC) ✍ Giovanni
- ✍ Calibration methods
- ✍ Cryogenics design ✍ Tom

4. Schedule

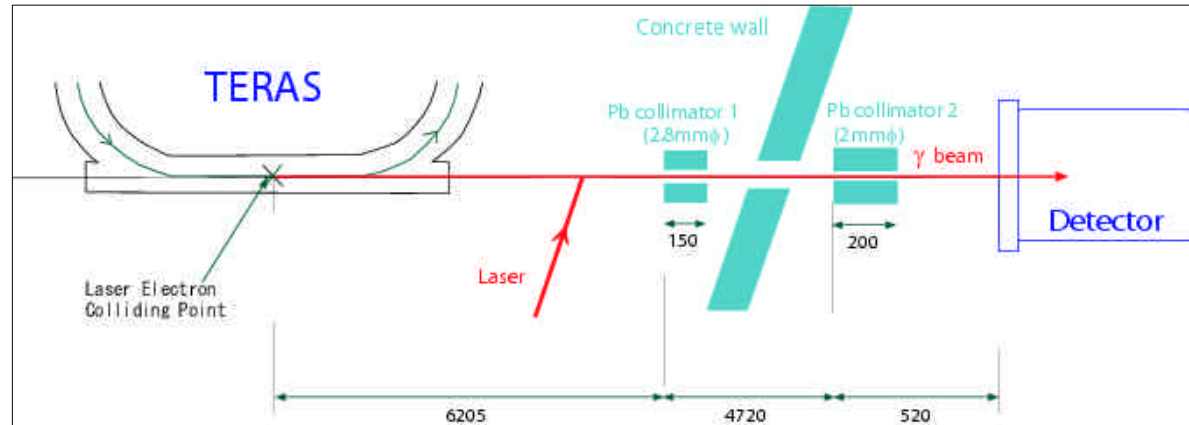


Gamma beam test

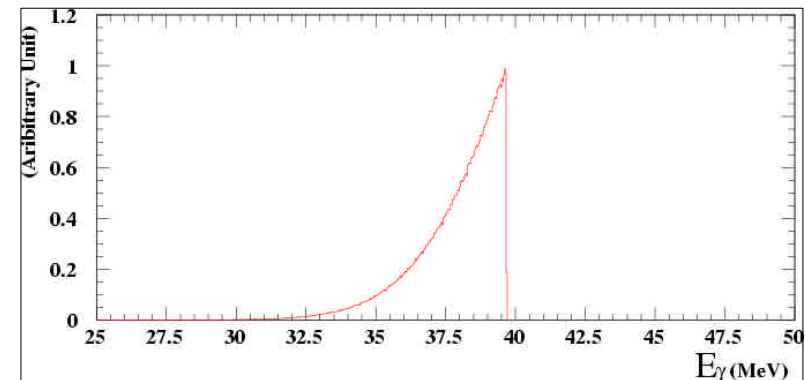
? beam at TERAS
Analysis



40MeV Compton Gamma at TERAS



- Electron Energy: 762 MeV.
- Max. current 200 mA.
- 40 MeV (20 MeV, and 10 MeV) Compton ? provided.
- Beam test in Feb. 2002 for 2 weeks

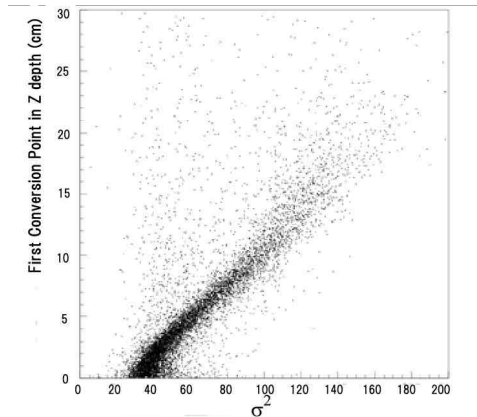




Energy Spectrum

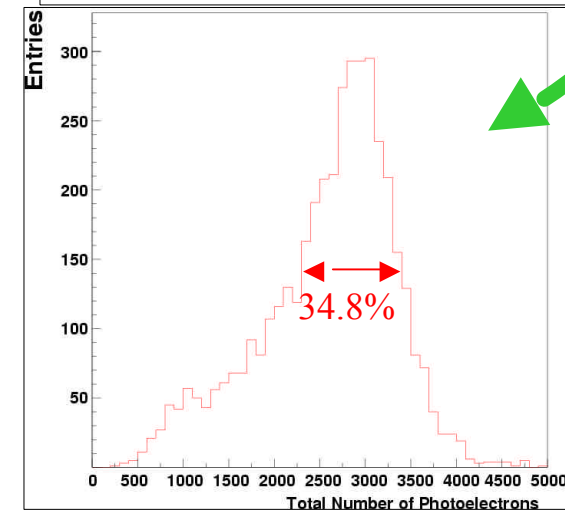
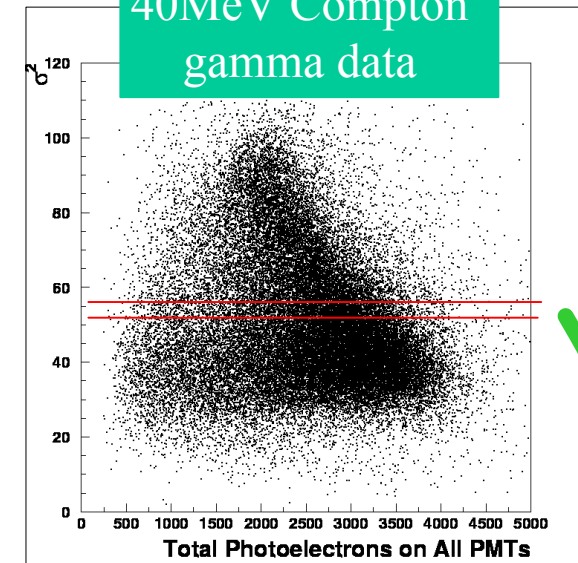
- σ^2 depth parameter:

$$\sigma^2 = \frac{\sum n_{pe}(i)x(i)^2 + n_{pe}(i)y(i)^2}{\sum n_{pe}(i)} - \left(\frac{\sum n_{pe}(i)x(i) + n_{pe}(i)y(i)}{\sum n_{pe}(i)} \right)^2$$



- Correlation between σ^2 and N_{pe}
~~short~~ σ^2_{abs} as explained later
- In the region of $50 < \sigma^2 < 55$
 34.8%(FWHM) including the energy spread of Compton ?

40MeV Compton
gamma data

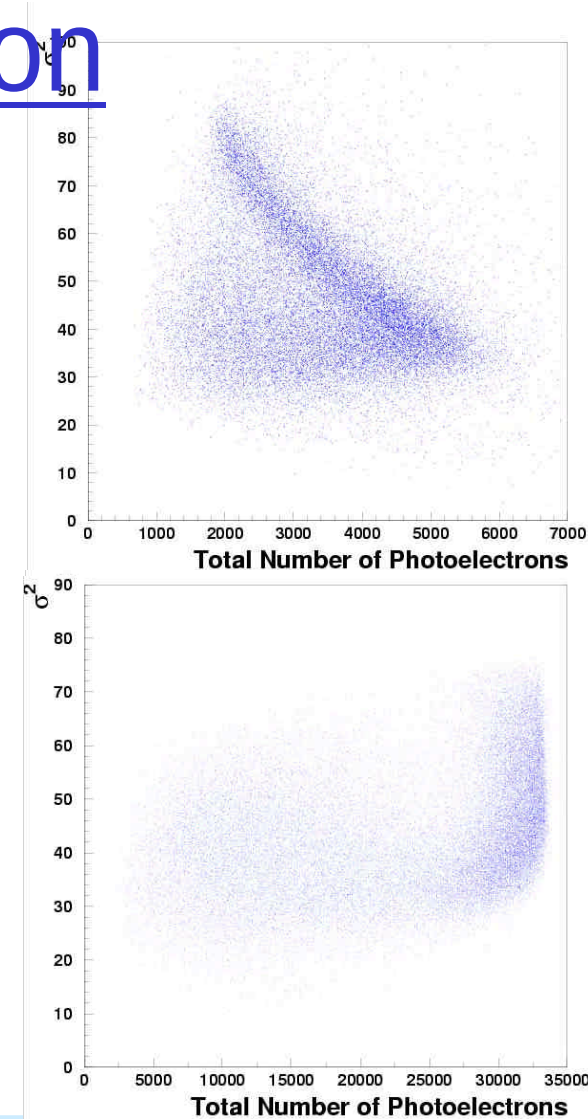


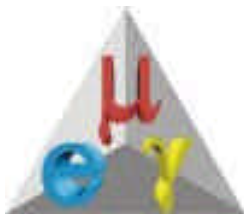
Satoshi Mihara for the $\mu \rightarrow e \gamma$ collaboration, review meeting at PSI, Jul 2002



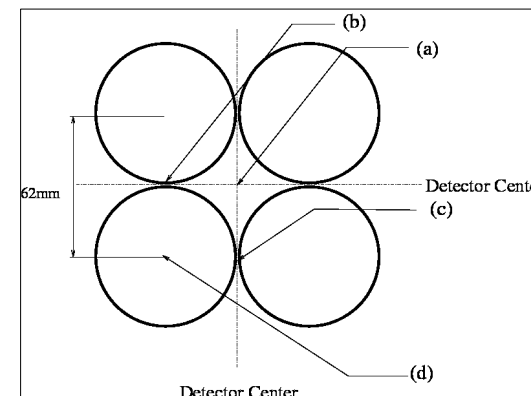
Comparison with MC with absorption

- Strong correlation between the σ^2 and N_{pe} can be explained by introducing absorption effect into MC.
- MC with 7cm σ_{abs} can explain the data.
- According to MC, σ_{abs} longer than 1m is necessary to achieve the resolution of a few % order.

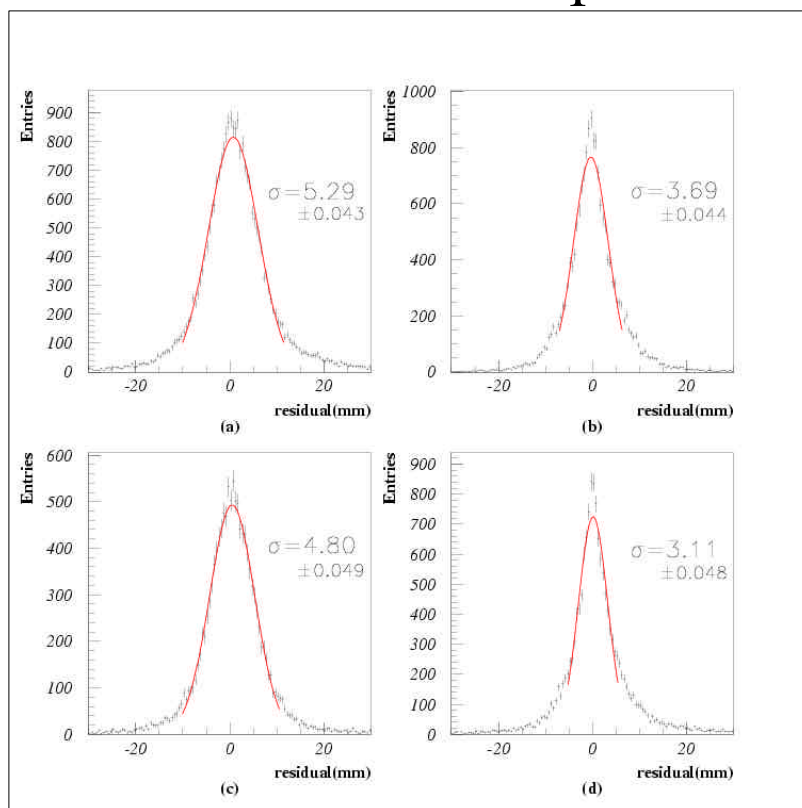




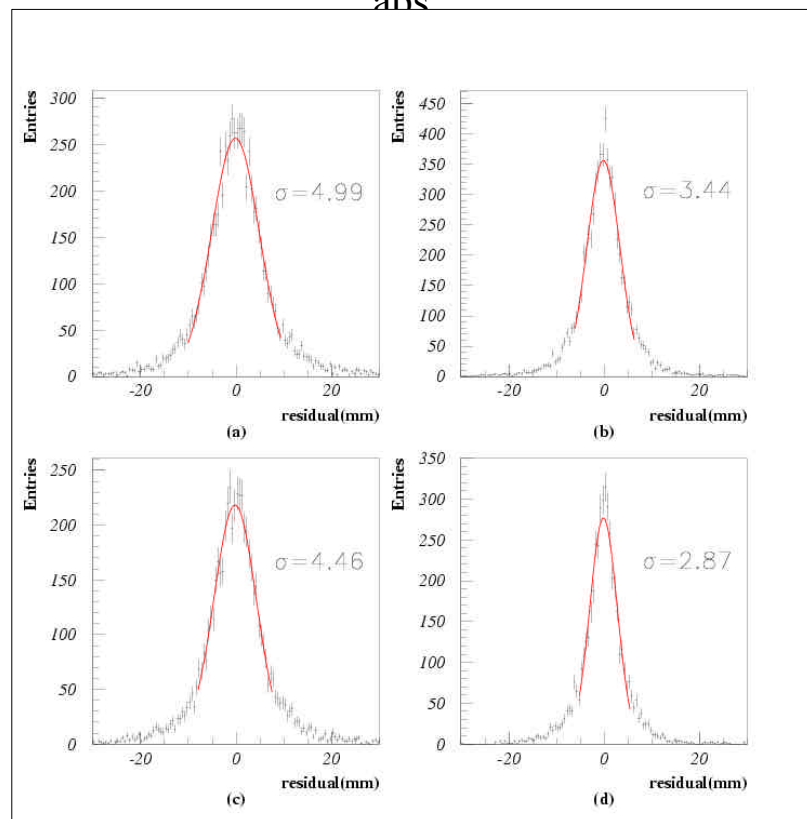
Position Resolution



- Data 40MeV Compton ?



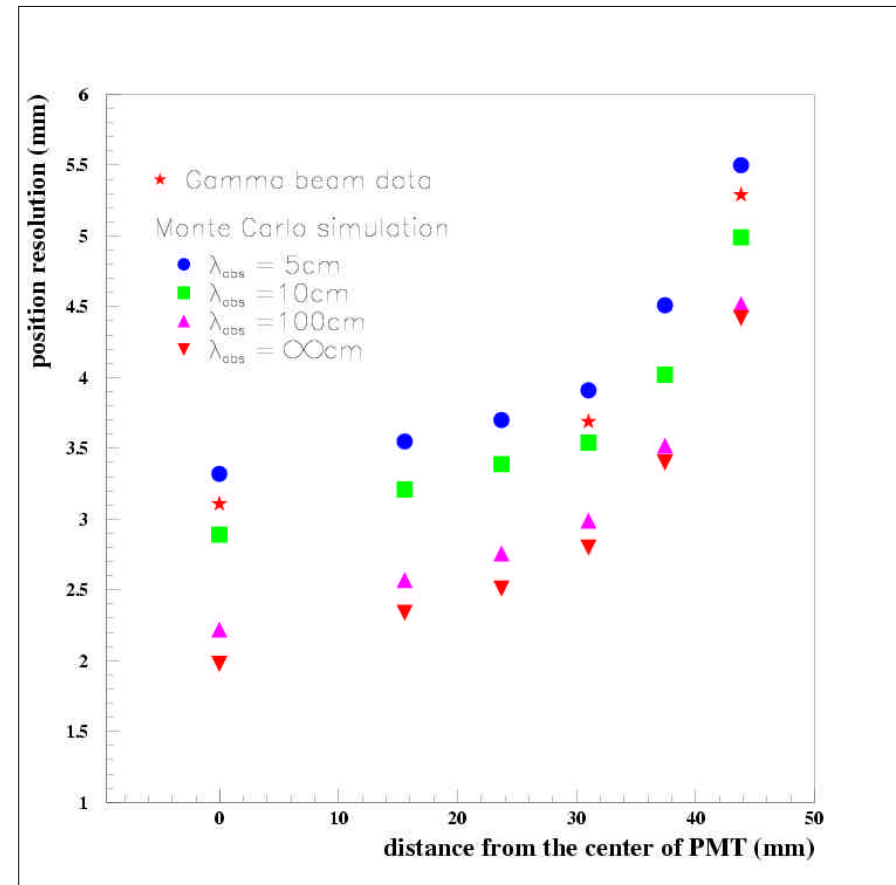
- MC with ?_{abs} = 10cm





Comparison with MC

- MC
 - $\lambda_{\text{abs}} = 5, 10, 100, 8 \text{ cm}$
 - $\lambda_{\text{abs}} = 30 \text{ cm}$
- Obtained resolutions agree with MC predictions including $5 < \lambda_{\text{abs}} < 10 \text{ cm}$.
- Further improvement expected with longer absorption length.





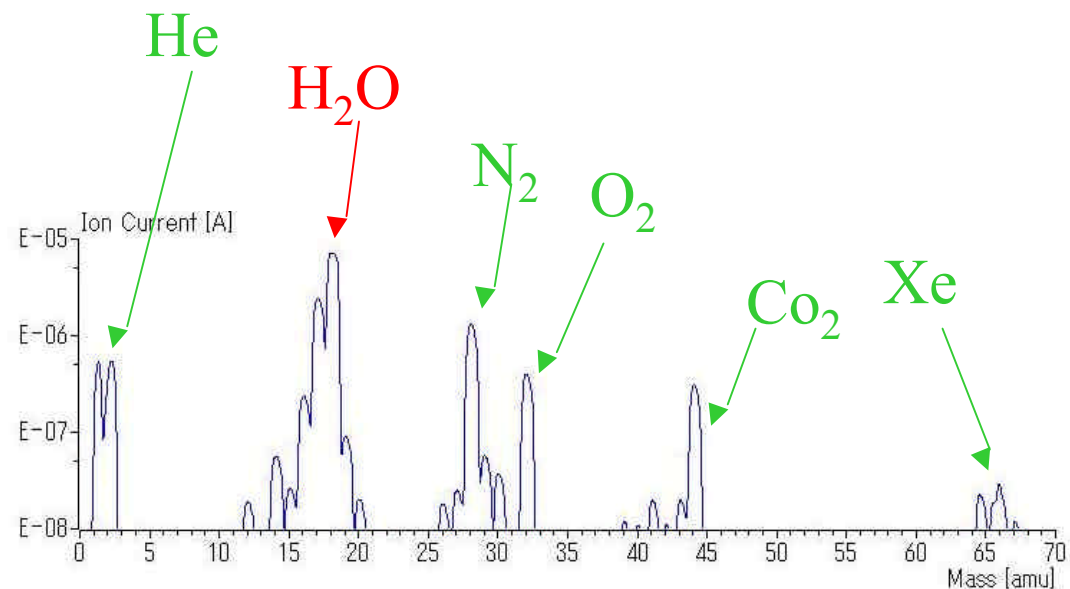
Xenon Purification



RGA with a mass spectrometer

- Remaining Gas Analysis (RGA) for investigating what causes short absorption length in gamma beam test.
- Remaining gas in the chamber was sampled to the analyzing section.

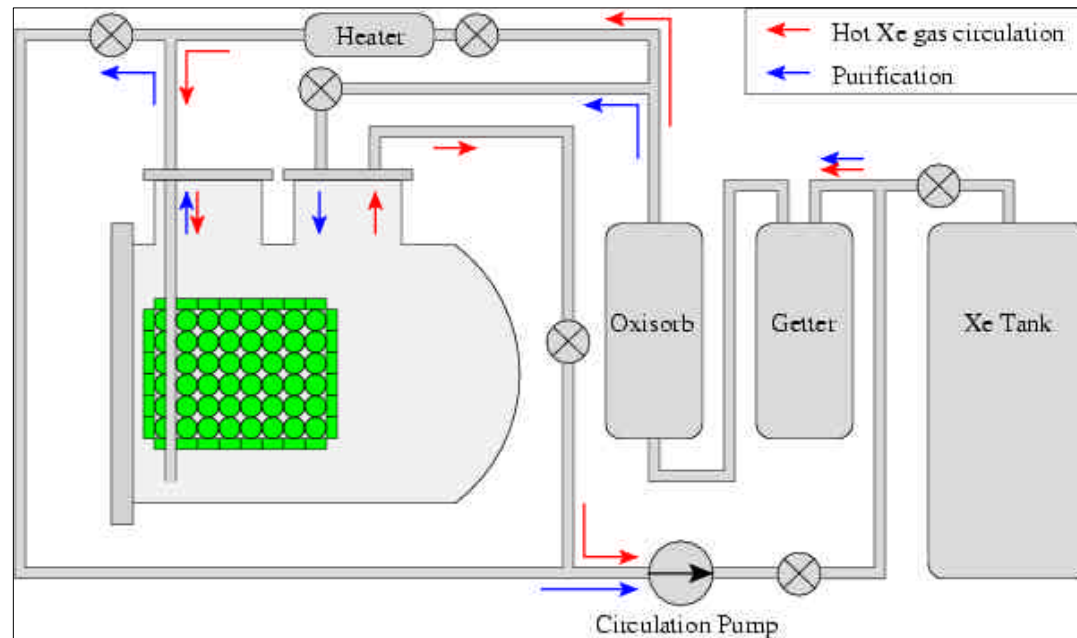
- Vacuum level
 - LP Chamber $2.0 \times 10^{-2} \text{Pa}$
 - Analyzing section $2.0 \times 10^{-3} \text{Pa}$





Purification System

- Xenon extracted from the chamber is purified by passing through the getter.
- Purified xenon is returned to the chamber and liquefied again.
- Circulation speed 10-12cc/minute

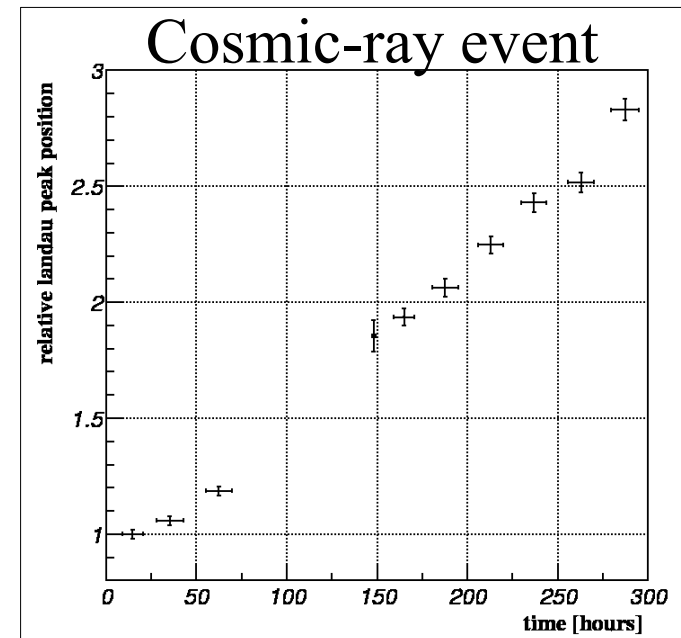
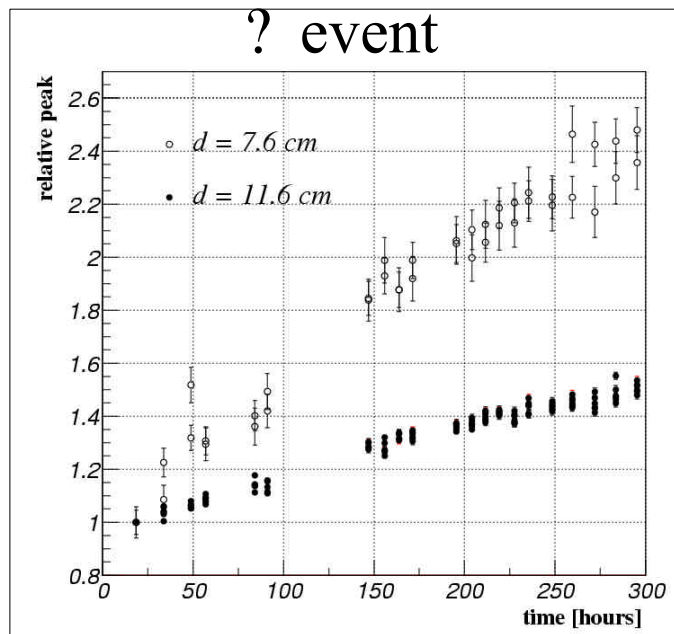


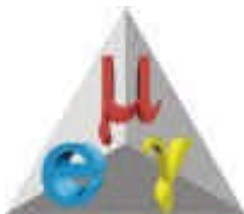


Performance of the purification

- Modeling:
$$N(r, t) = N_{\infty} \exp\left(\frac{r}{\lambda_0} \exp\left(\frac{t}{\tau}\right)\right)$$

if no continuous outgassing

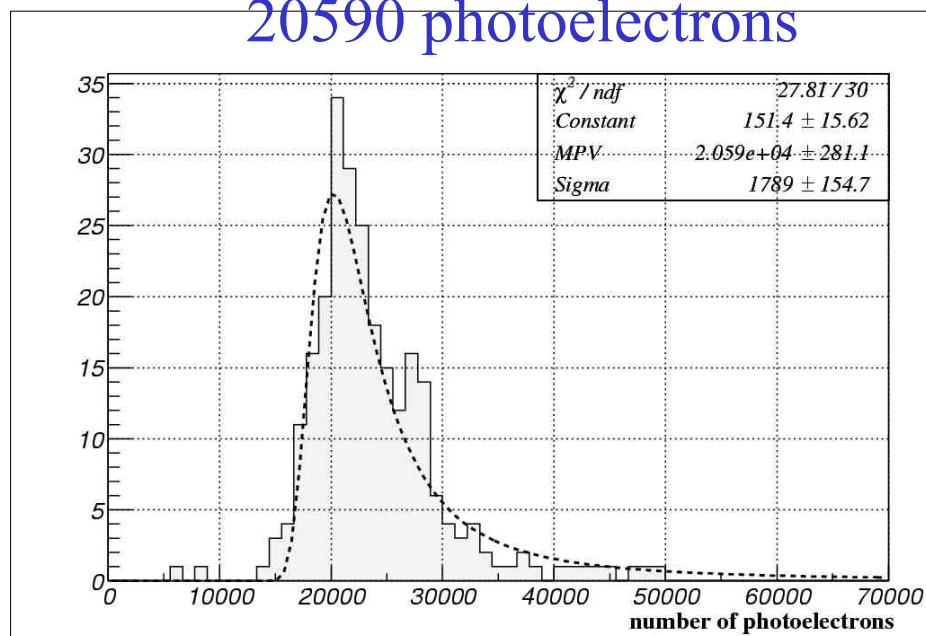




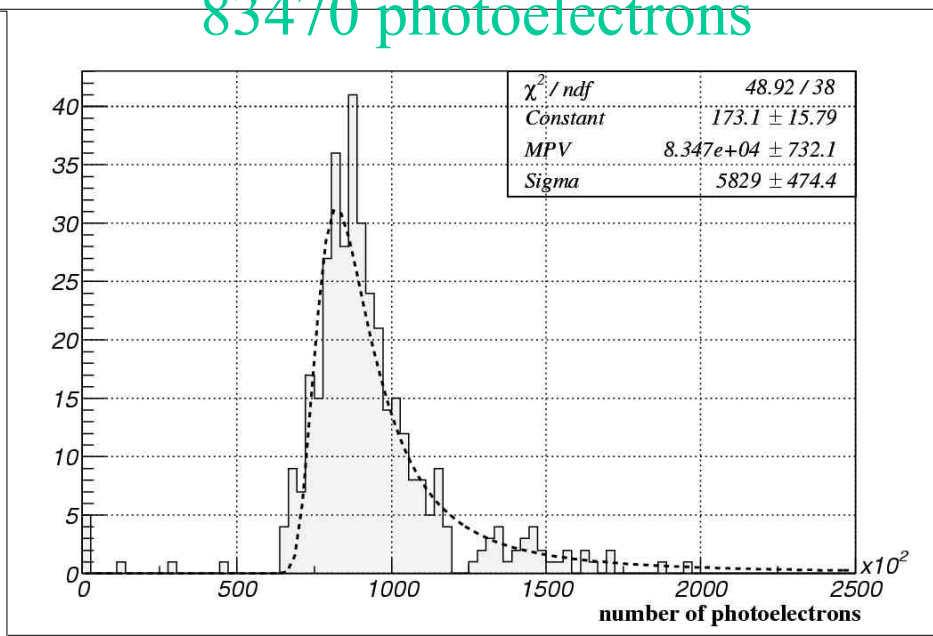
After 600 hours...

- Light yield increased by factor of 4.
- Comparison with MC prediction $\not\approx$?_{abs} > 1m

Before purification
20590 photoelectrons



After purification
83470 photoelectrons





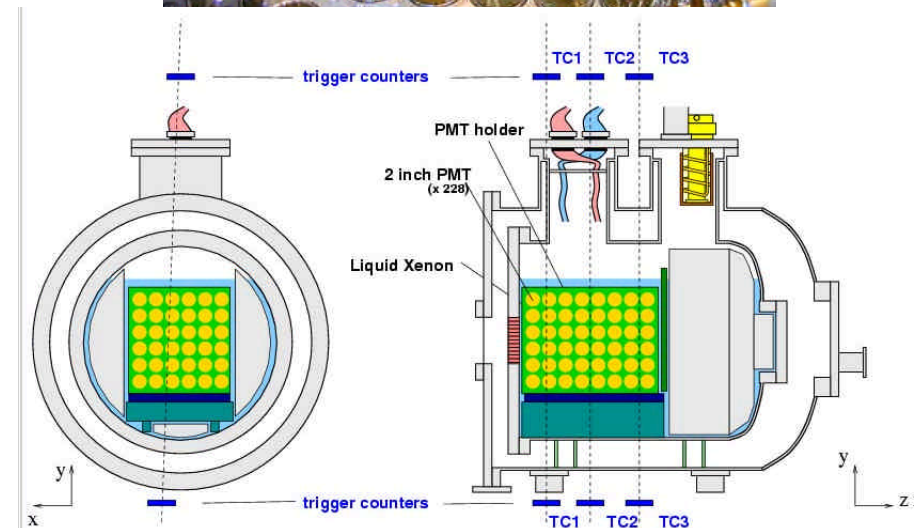
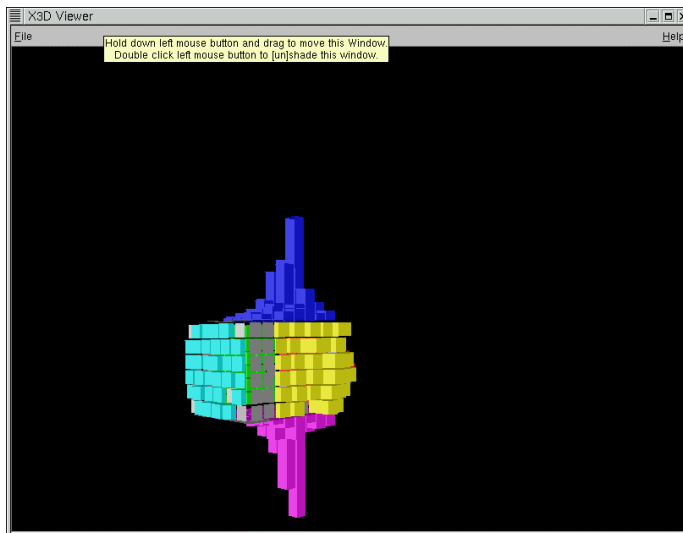
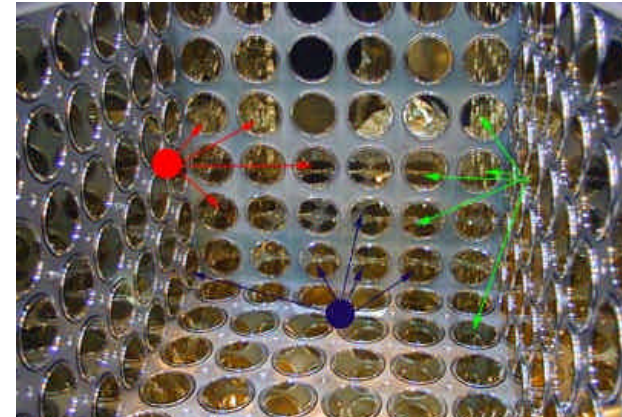
Absorption Length Measurements

before purification
after purification



Estimation of Absorption Length

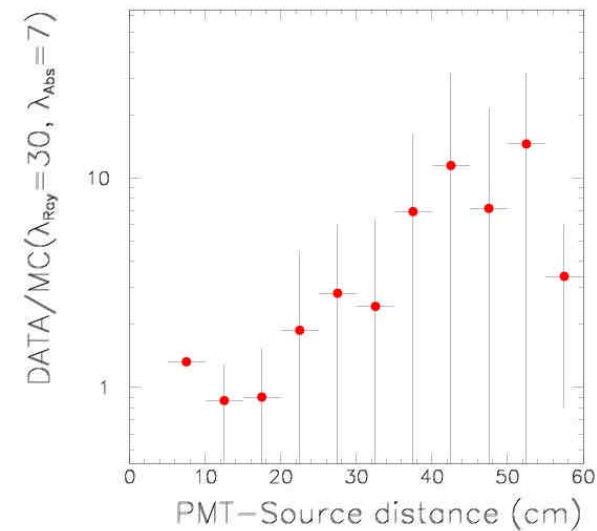
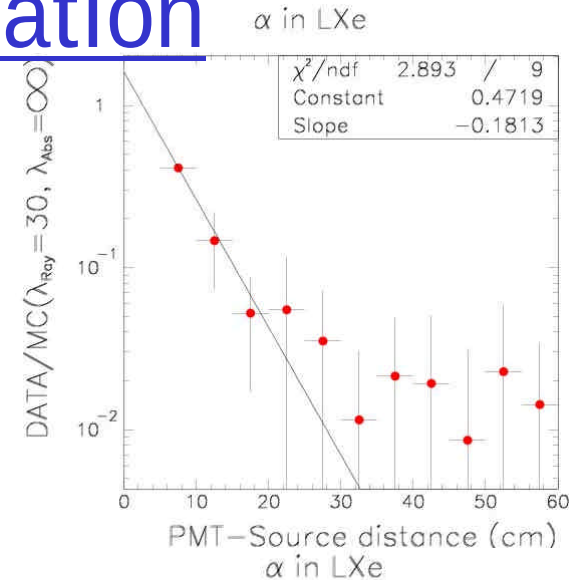
- ? & cosmic-events.
 - Relation between the light yield and distance from the light source to PMTs.





Absorption Length (?) before purification

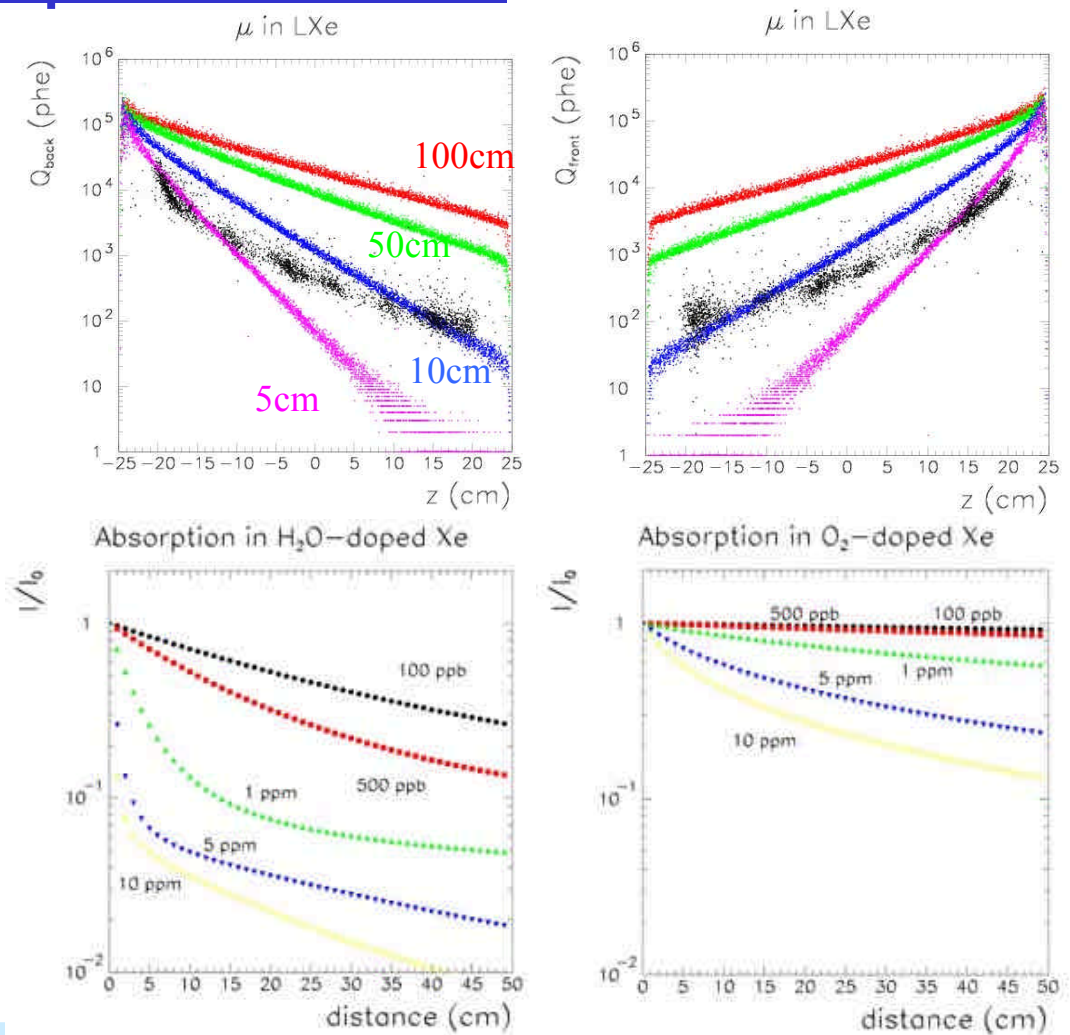
- Data/MC as a function of the distance from the alpha source to the PMTs
 - MC: $\lambda_{\text{Ray}} = 30\text{cm}$
 $\lambda_{\text{abs}} = 8 \text{ cm and } 7\text{cm}$
- MC with 7cm absorption can explain rapid decrease of the light yield at short distance.





Absorption Length (CR) before purification

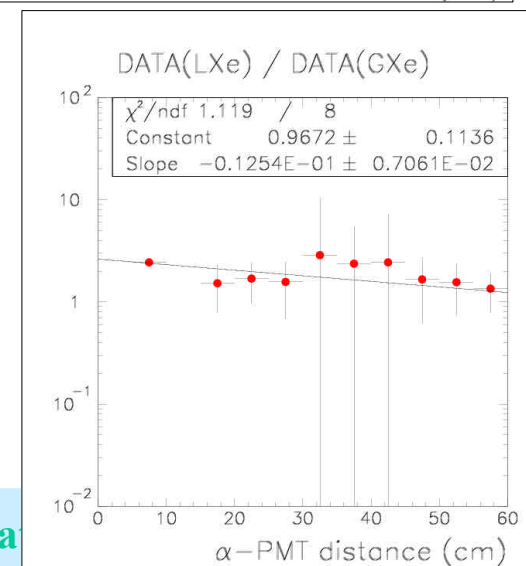
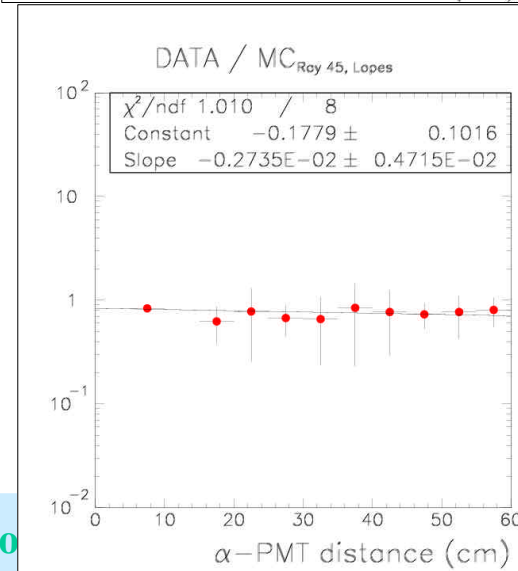
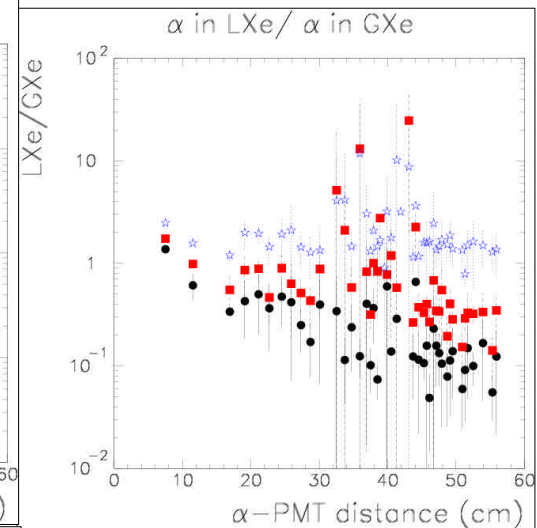
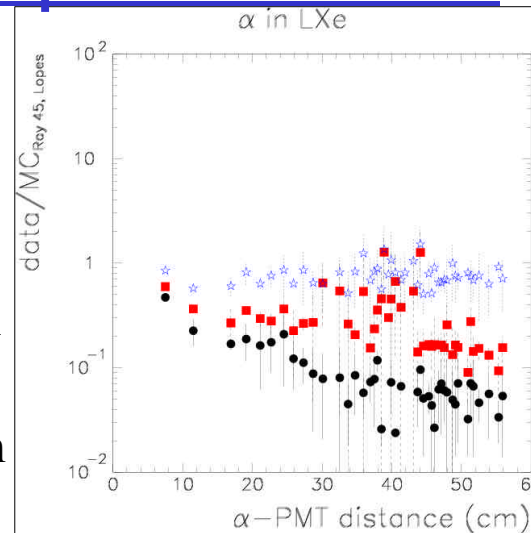
- Data distribution is steeper near the face and falls less violently for large z .
- The discrepancy can be explained by introducing wavelength-dependent absorption effect by water.
- Absorption length: 5~10cm



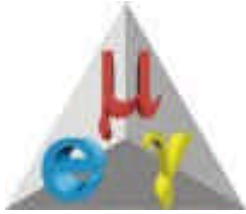


Absorption Length after purification

- Fit the data with a function :
 $A \exp(-x / \lambda_{\text{abs}})$
- $\lambda_{\text{abs}} > 100\text{cm}$ (95% C.L) from comparison with MC.
- ($\lambda_{\text{abs}} > 80\text{cm}$ from comparison with cold gas data which however includes diffusion effect)
- CR data indicate that $\lambda_{\text{abs}} > 100\text{cm}$ has been achieved after purification.

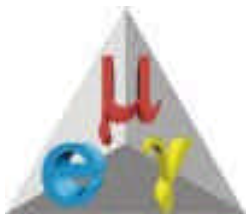


Satoshi Mihara for the $\mu \rightarrow e \gamma$ co



Other related R&D

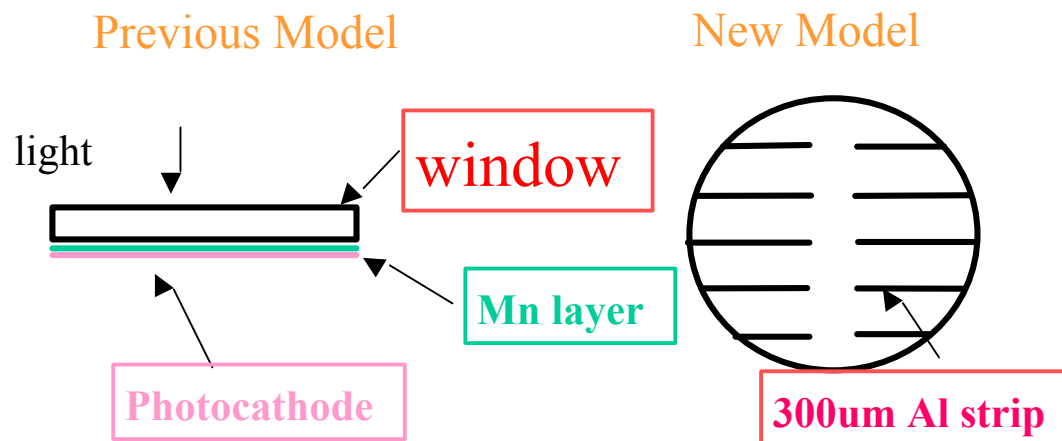
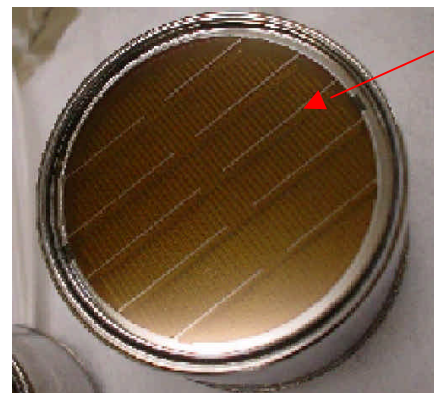
PMT
Purity monitor



PMT Development

- The previous model used a Mn layer to keep the surface conductivity of the photocathode at low temperature.
- The new model uses Al strip instead of the Mn layer.
- QE is expected to improve and PMT production in more constant quality.

Aluminum Strip

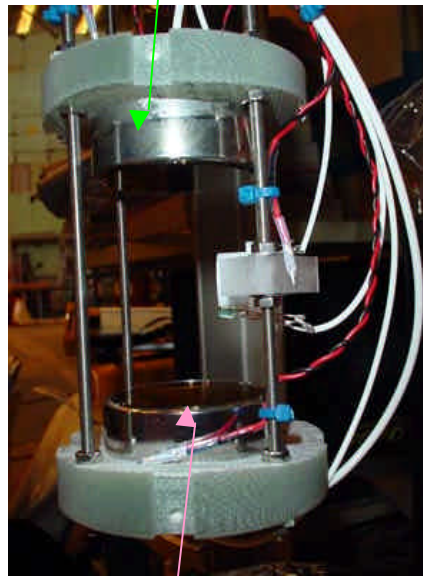




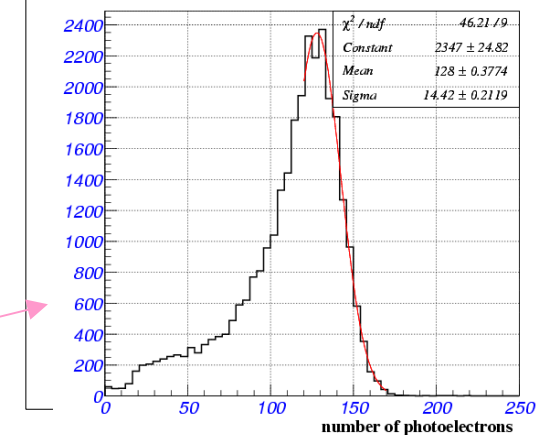
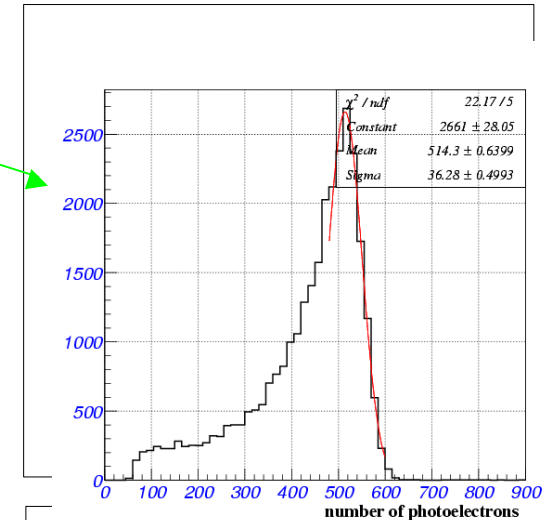
PMT test in gas/liquid xenon

- Tests in cold gas and in liquid xenon performed.
- QE improved by factor of 2-3.
- Rate dependence is similar or slightly better. (need careful check)
- K-Cs-Sb photocathode can probably be used (Previous model with Rb-Cs-Sb) ✗ possible to achieve higher gain with same HV.

New model



Previous model

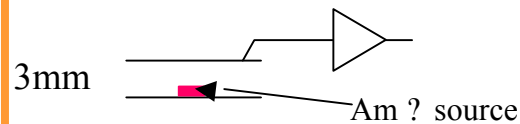




Purity Monitor

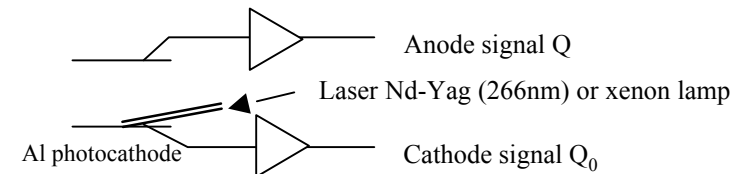
1. By measuring concentration of electro-negative impurities

- ? source ionization chamber



- Simple and stable.
- Possible to measure impurity of 10~100ppb.
- Signal amplitude is rather small, a few mV level for 100ppb impurity.
- Implemented in LP.

- Laser-induced ionization chamber



- Laser stability can be monitored by measuring cathode signal.
- Large amount of light yield.
- Possible to measure impurity 1ppm~1ppb.
- Need development, but we can use a similar one as ICARUS developed.

2. By measuring absorption of scintillation light

- | | |
|-----------------------|--|
| • Cosmic-ray event | Possible to measure $?_{abs} > 1m$ but low rate. |
| • ? event | Possible to measure $?_{abs} \sim 1m$ but small signal. |
| • Other possibilities | Direct observation of laser light through xenon (Exima laser is a candidate, under investigation). |

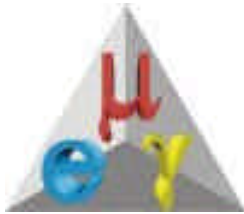


Final Detector

calibration

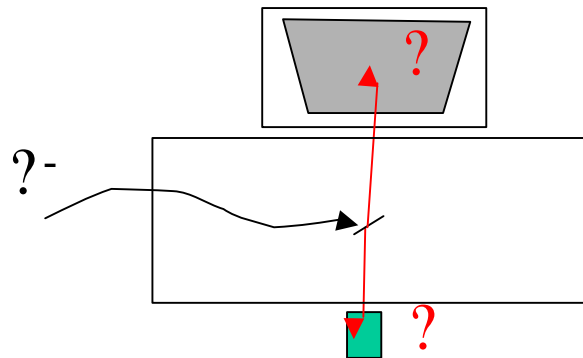
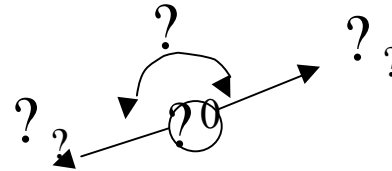
expected performance

cryogenics design



Calibration

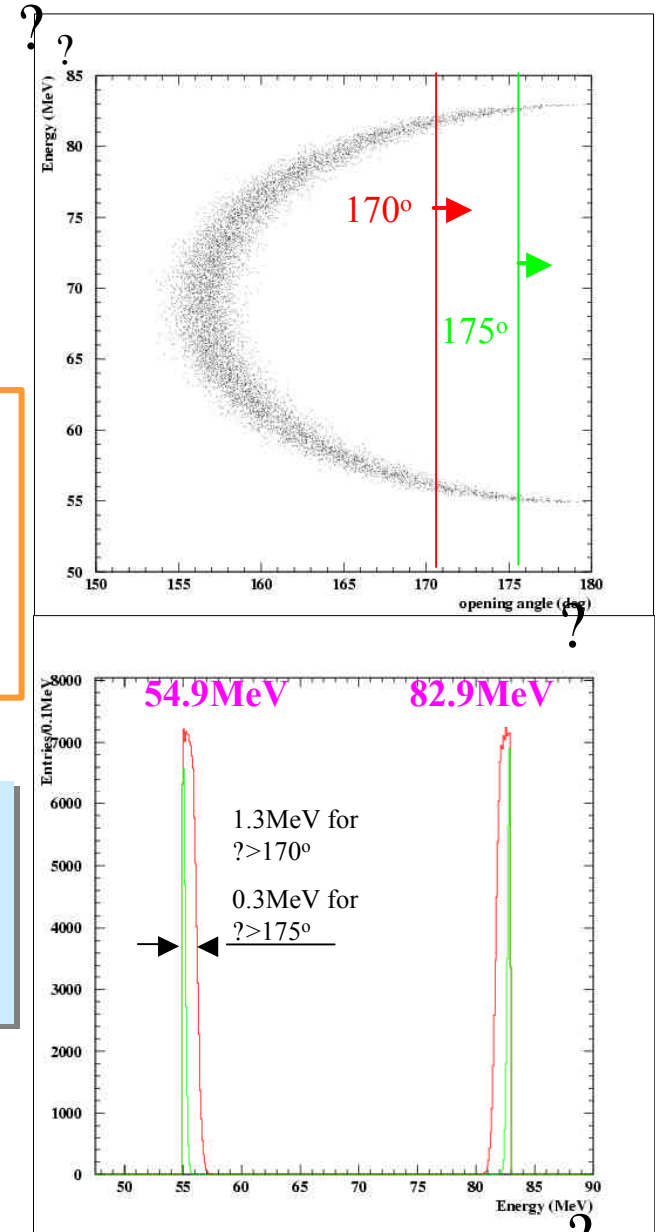
?? γ - $p \rightarrow n$
 $\gamma^0(28\text{MeV}/c) \rightarrow ??$
 $??\text{MeV} ??\text{ eV}$

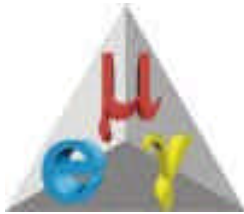


- Requiring $???$
FWHM = 1.3 MeV
- Requiring $? > 175^\circ$
FWHM = 0.3 MeV

- No need of excellent energy resolution.
- Position resolution of $? = 4\sim 5\text{cm}$ is enough.
- Timing resolution ($? < \sim 1\text{nsec}$) is required for timing calibration of the xenon detector.

2. Am-Be γ source 4.43 MeV

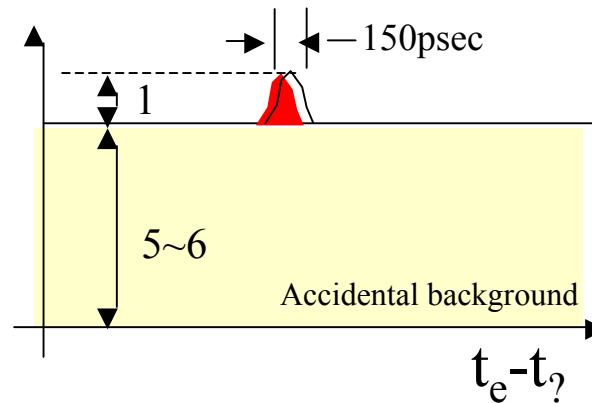




Calibration cont'd

- $? \rightarrow e \gamma \gamma$
- $E_e > 0.85 E_\gamma > 0.8 \quad ?_{e\gamma} > 120^\circ$

$10^8 ? / \text{sec}$



$10^7 ? / \text{sec}$

Signal 1/10
Background < 1/100

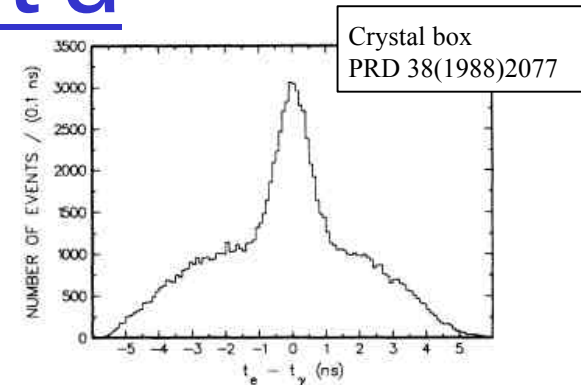
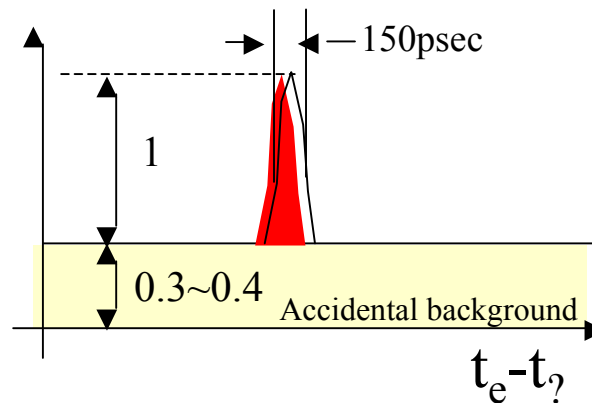
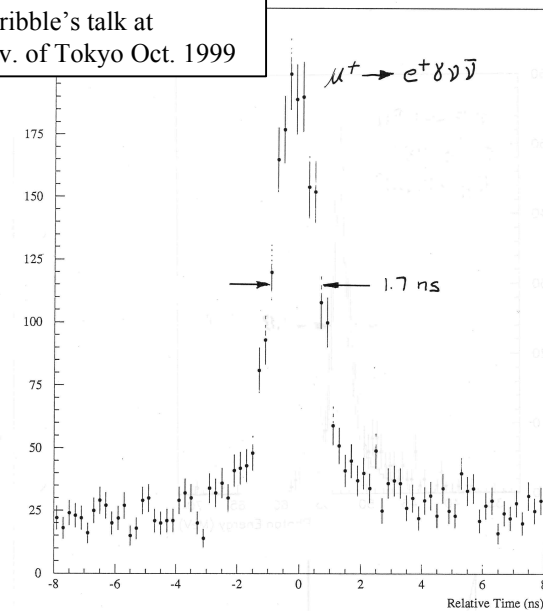


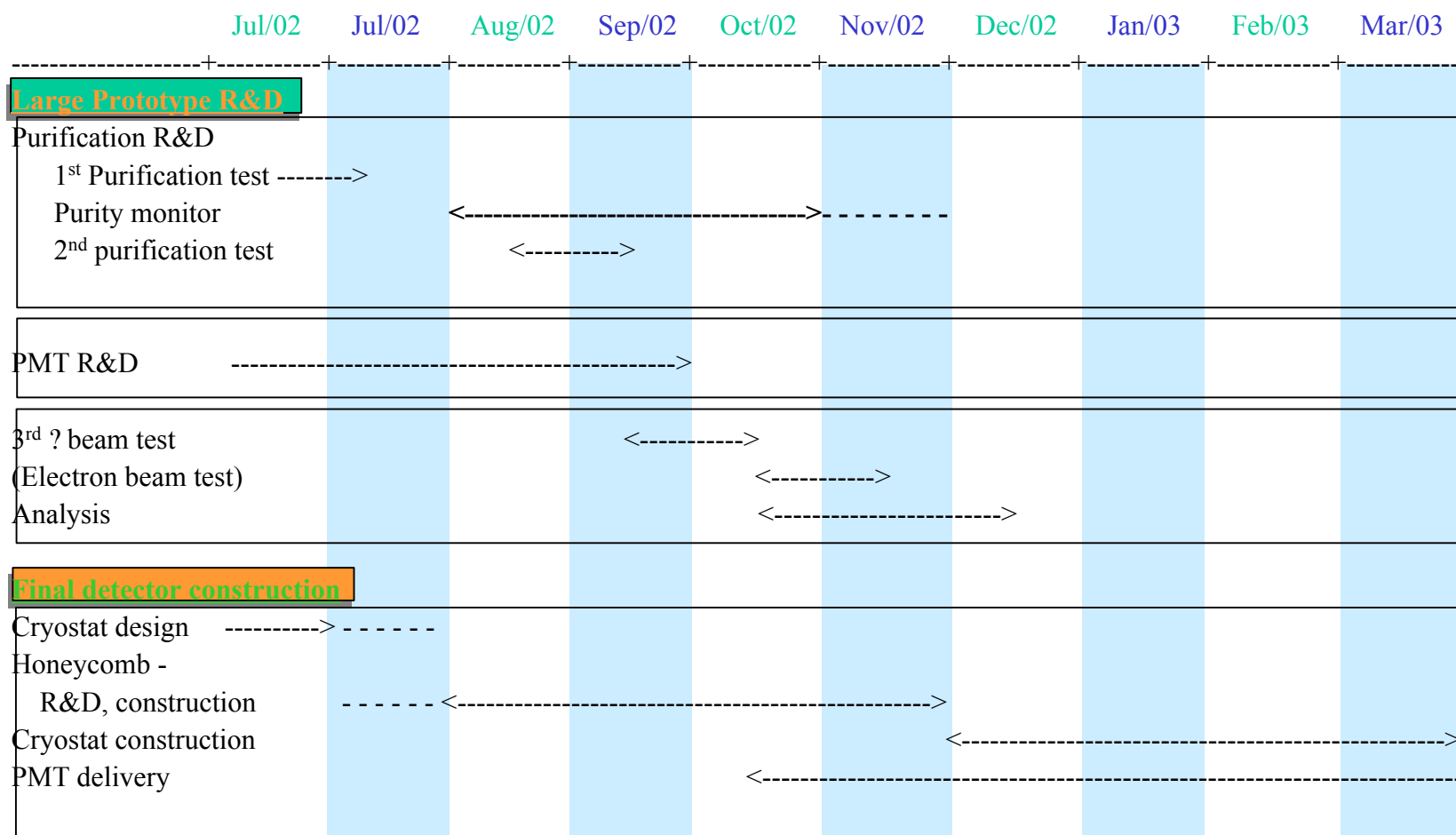
FIG. 9. The relative positron-photon timing for $e\gamma$ triggers.

R. Tribble's talk at
Univ. of Tokyo Oct. 1999





Schedule





Summary

- 2nd ? beam test in Feb. 02
 - Worse resolutions than our expectation due to short absorption length caused by contaminant impurity.
- Purification system has been developed and 1st test was successfully done.
 - Recent CR and alpha data indicate $\tau_{\text{abs}} > 1\text{m}$.
 - Increasing purification speed is the next step for quick start of the detector operation.
 - Development of purification monitor is an important issue.
- Another tests are planed with purified xenon using ? and e beams in autumn.
- Other R&D works are going.